

**Ray tracing and Monte Carlo methods for
uncertainty quantification in the pointing
direction of radio telescopes for CMB
observations.**

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18 Luglio 2025

Motivations for this thesis

Cosmology and The Big Bang Theory

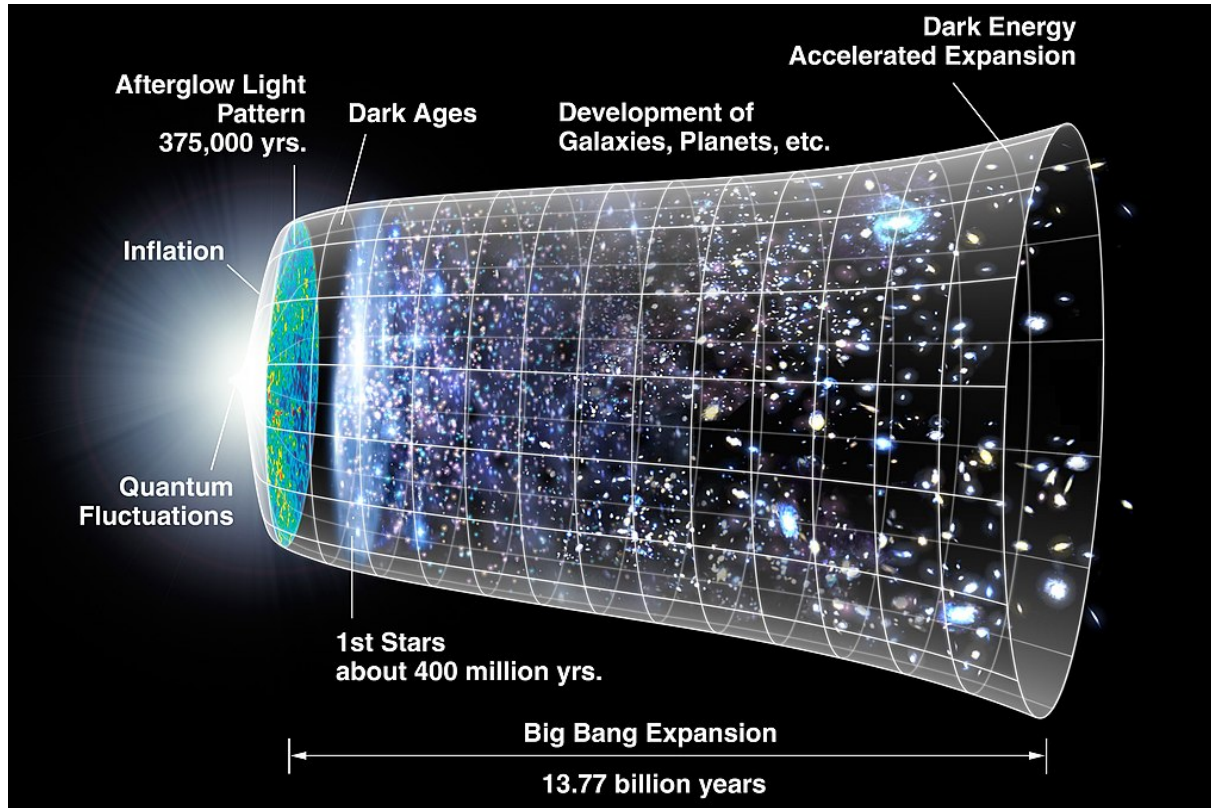


Figure 1: Evolution of the Universe according to the Big Bang theory and inflation.

Cosmic Microwave Background

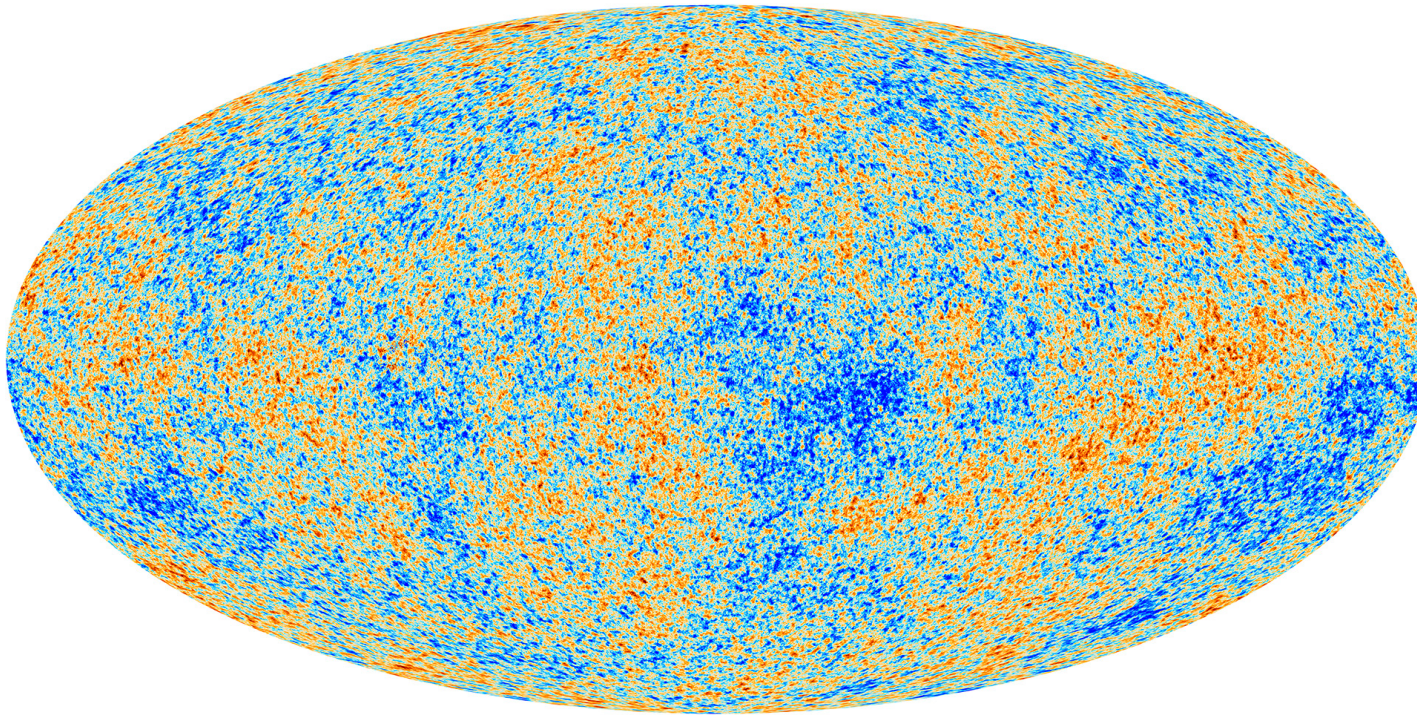


Figure 2: CMB observed by Planck instrument, image from ESA.

Motivations for this thesis

CMB Radio Telescopes

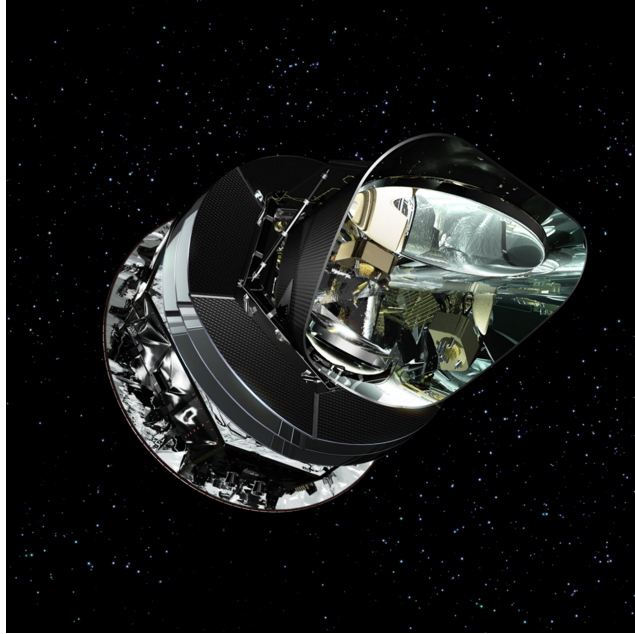


Figure 3: The Planck spacecraft, image from ESA.

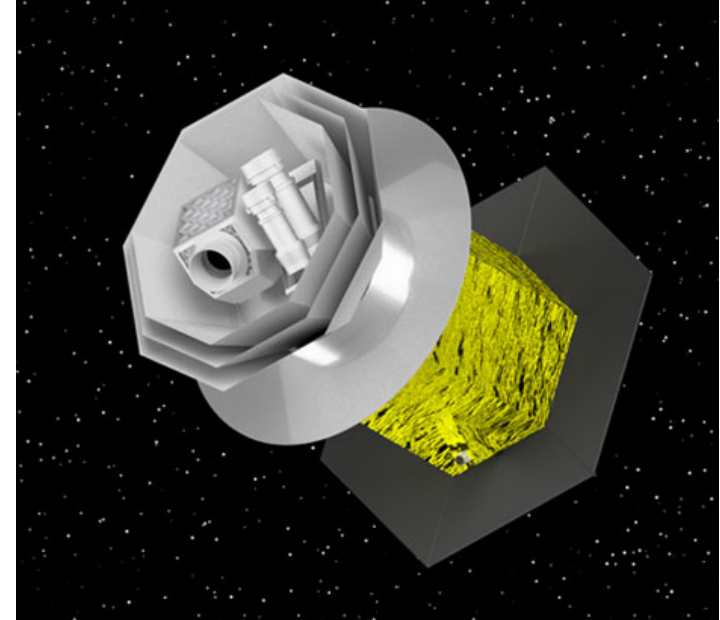


Figure 4: The LiteBIRD spacecraft, image from ISAS.

Motivations for this thesis

CMB Radio Telescopes

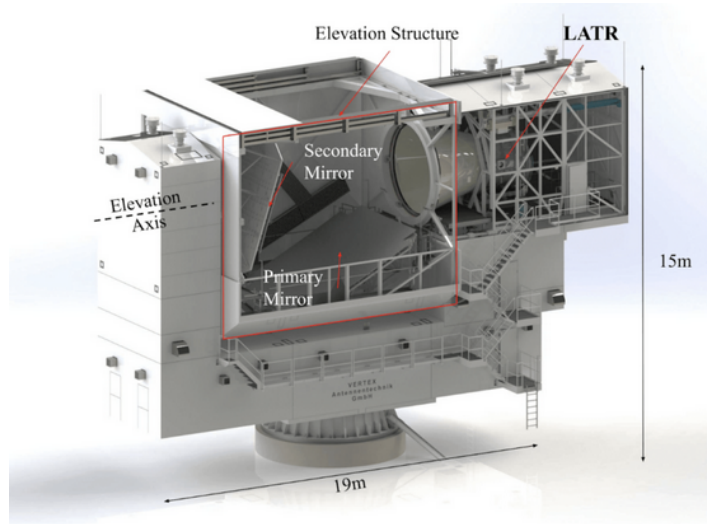


Figure 5: LATR telescope at Simons Observatory, in the Atacama Desert of Northern Chile.

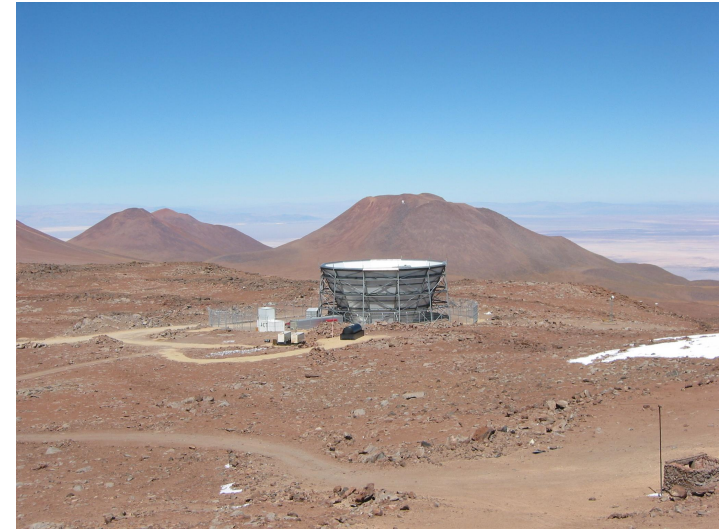


Figure 6: Atacama Cosmology Telescope (ACT), in the Atacama Desert of Northern Chile.

The Goal

- The goal is to investigate how **mechanical perturbations** in the **optical system** of a **radio telescope** influence the measurements of the **Cosmic Microwave Background**.

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- There are already **proprietary softwares** (e.g., TICRA) that offer both ray-tracing and more accurate full-wave analysis, but they are computationally and **economically expensive**.
- In many cases we can just rely on **geometrical optics**, or often called **ray optics**.

Ray Optics

- Fermat's **Principle of Least Time**
→ laws of **Reflection** and **Refraction**.

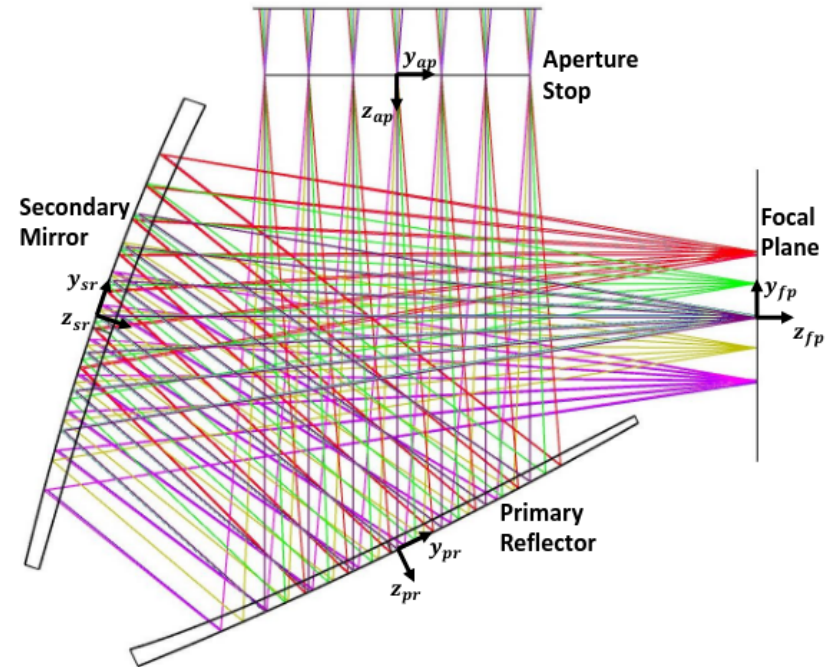
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Ray Optics

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→ optical path taken by light is **reversible**
→ trace rays from the focal plane to the sky



Ray-Tracing + Monte Carlo

- Using the **Julia** programming language, we created a **software** that can model **any telescope's reflective optical system** and implements a ray-tracing algorithm to assess the **pointing direction**.
- Jointly with **Monte Carlo simulations**, it is possible to explore **many configurations** of the system under a vast range of perturbations, in order to **propagate the uncertainty** in the telescope configuration into the measurements.

Motivations for this thesis

The LSPE/STRIP telescope

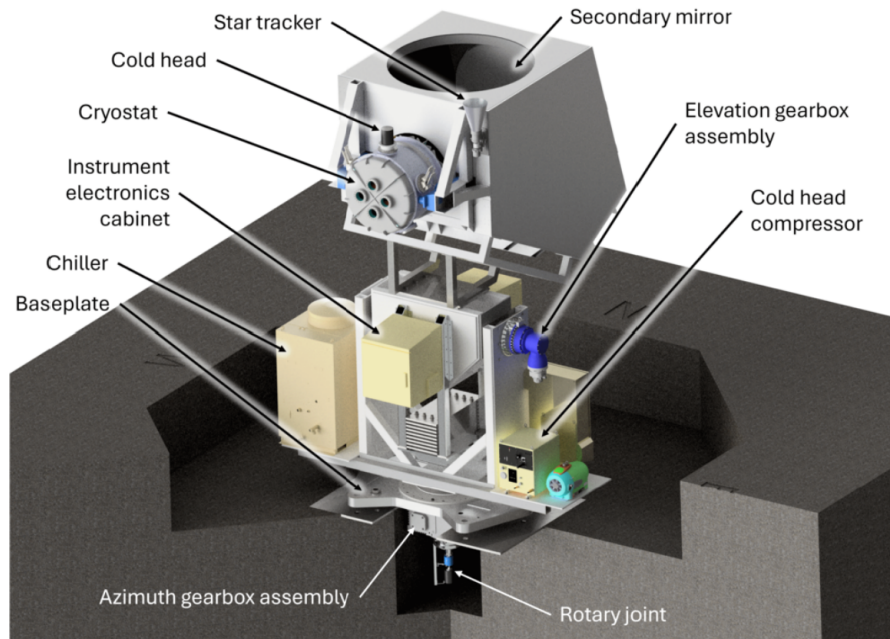


Figure 7: Rendering of LSPE-STRIP full instrument.

- As a case study, we investigated the pointing accuracy of the *LSPE/STRIP*'s dual-reflector crossed-Dragone optical system under **tilting perturbations**.

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The LSPE/STRIP telescope

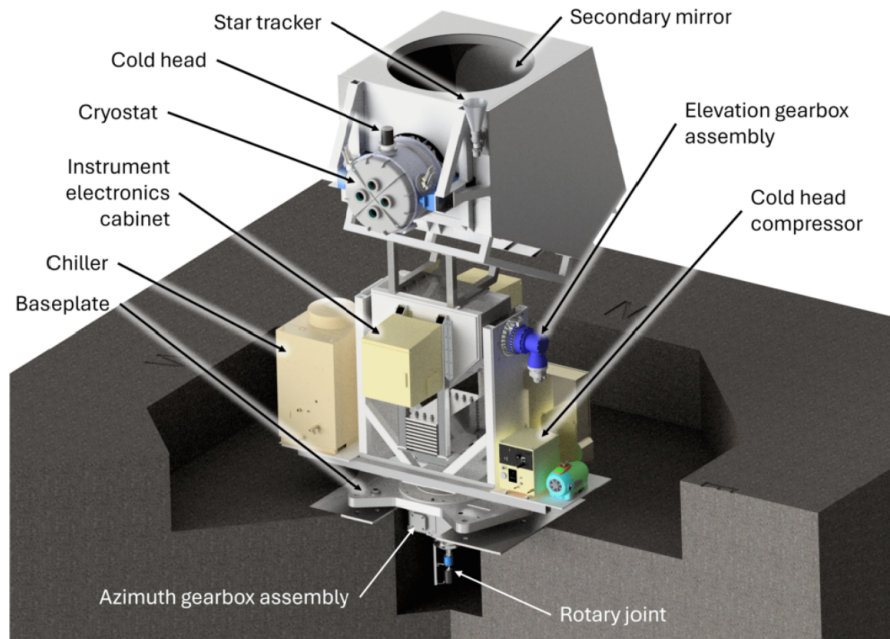


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- As a case study, we investigated the pointing accuracy of the *LSPE/STRIP*'s dual-reflector crossed-Dragone optical system under **tilting perturbations**.
- We **integrated** the pointing direction, evaluated with the ray-tracing software, directly into the STRIP **end-to-end pipeline** to simulate sky observations.

Ray-Tracing & MC Simulations

Tilted Reflectors and Pointing Error

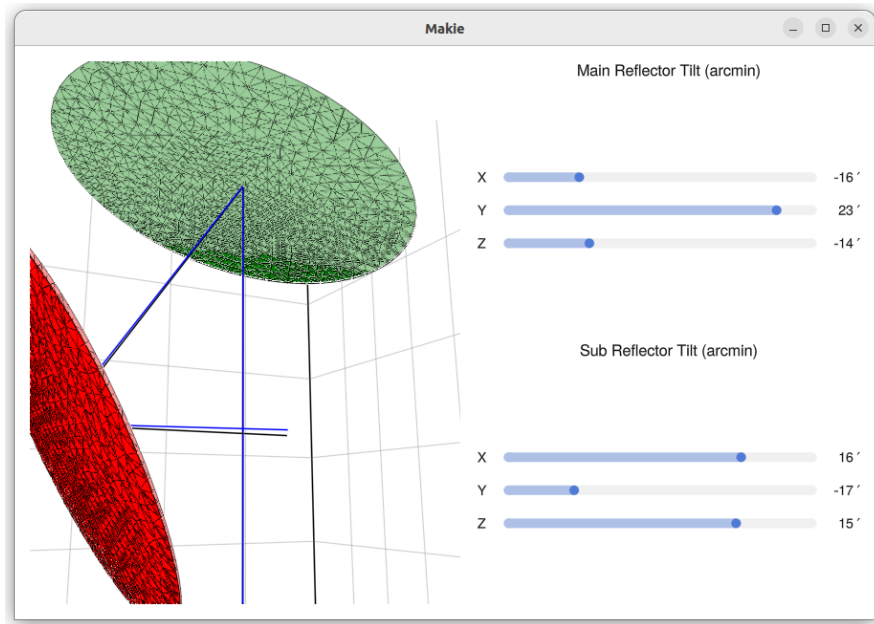


Figure 8: Interactive demo of a tilting simulation implemented with our ray-tracing engine and `Makie.jl`.

Tilted Reflectors and Pointing Error

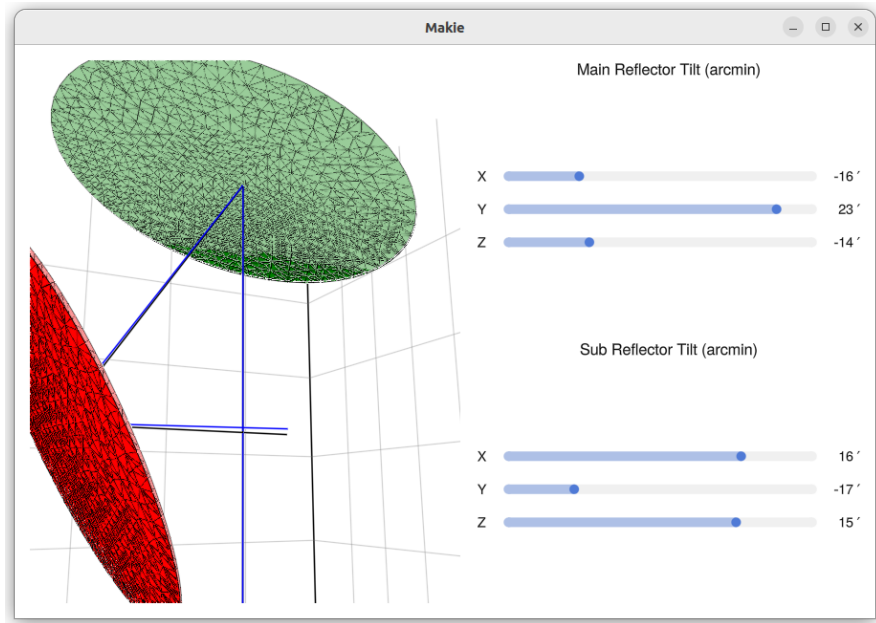


Figure 8: Interactive demo of a tilting simulation implemented with our ray-tracing engine and `Makie.jl`.

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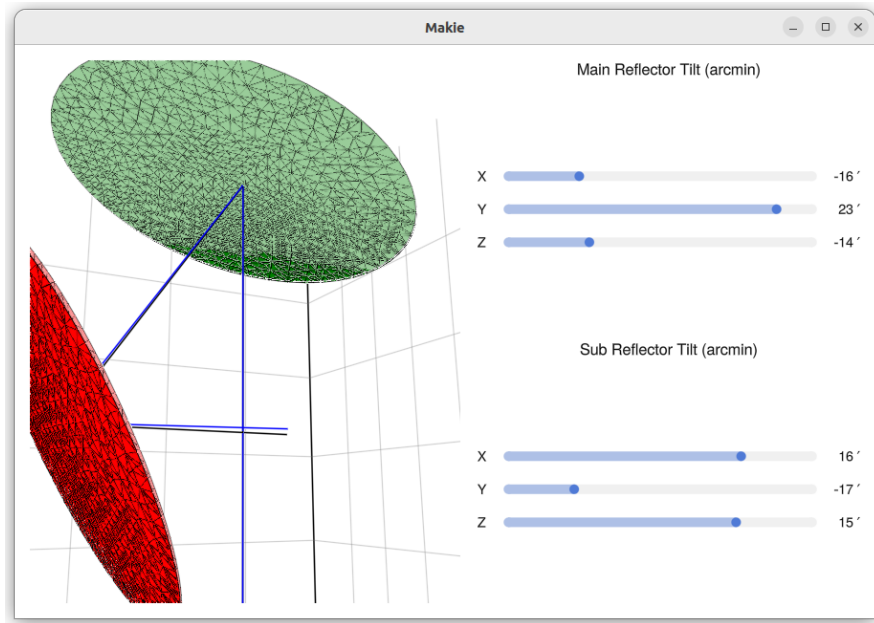


Figure 8: Interactive demo of a tilting simulation implemented with our ray-tracing engine and `Makie.jl`.

- We first create the **ideal configuration** of the optical system and trace a ray from the focal plane of the instrument to get the **ideal pointing direction** in the sky.
- Then we **perturb** the ideal system and repeat the process, obtaining a **tilted direction**.

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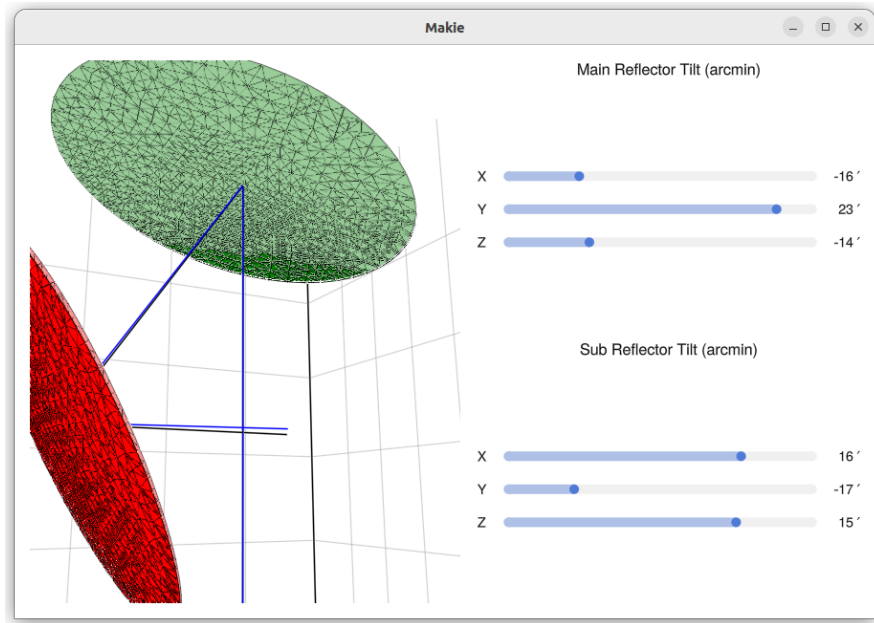


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$$\text{POINTING ERROR} \equiv \arccos(\hat{P}_{\text{ideal}} \cdot \hat{P}_{\text{tilted}})$$

Ray-Tracing & MC Simulations

Simulations SetUp

BCV
progetti s.r.l.

INFN – sez. Milano
LSPE experiment
STRIP telescope

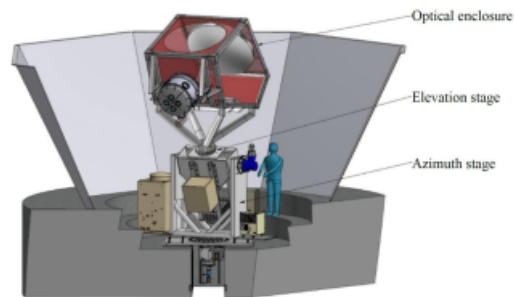
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tel. (+39)-0286-452002

Large Scale Polarization Explorer (LSPE) STRIP telescope

TECHNICAL NOTE #7

WP2 – AZIMUTH STRUCTURES – REVISED DESIGN,
FULL TELESCOPE ANALYSIS



- We consider **independent** tilt angles

$$\theta_{a_i} \sim \mathcal{N}(0, \sigma^2)$$

$$a \in \{\text{primary, secondary}\}, \quad i \in \{x, y, z\}$$

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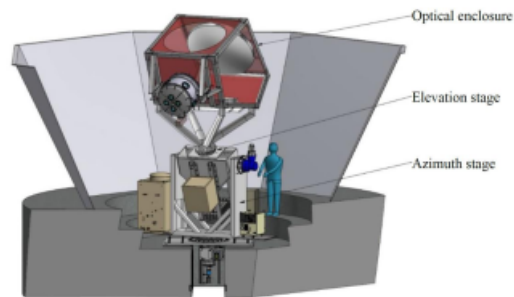
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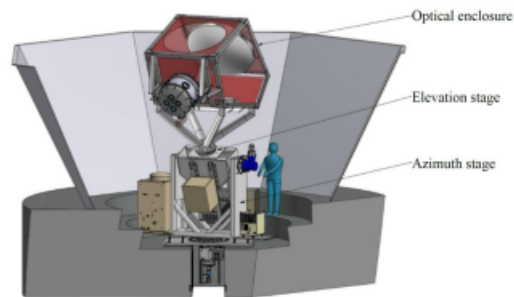
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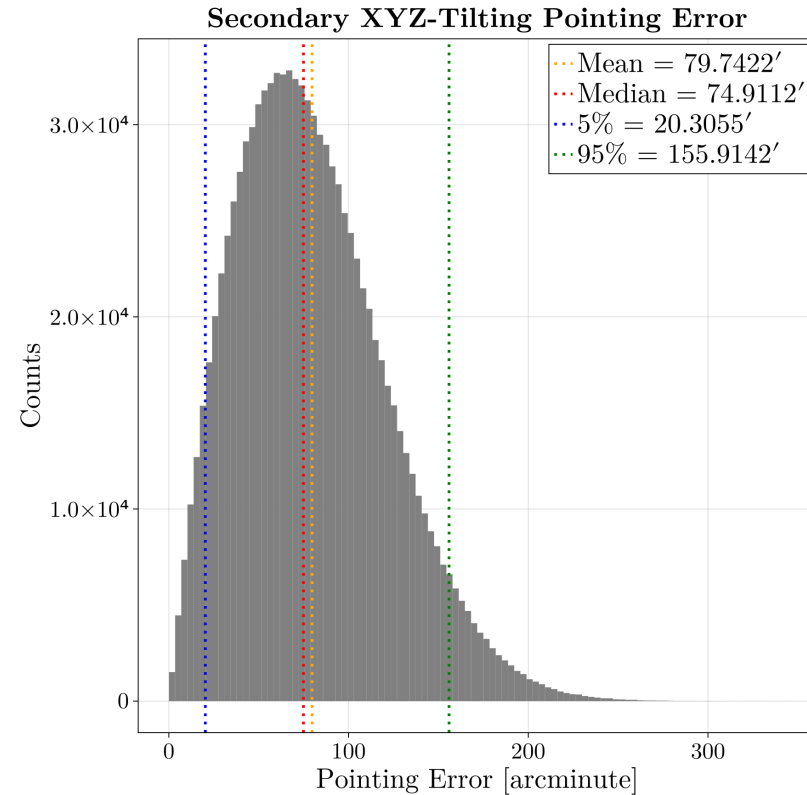
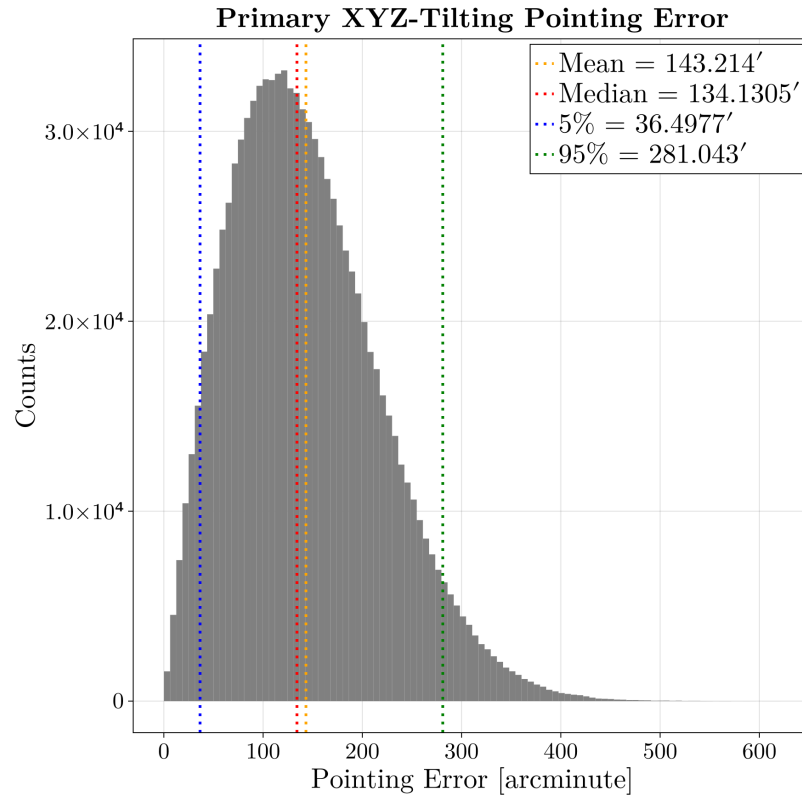
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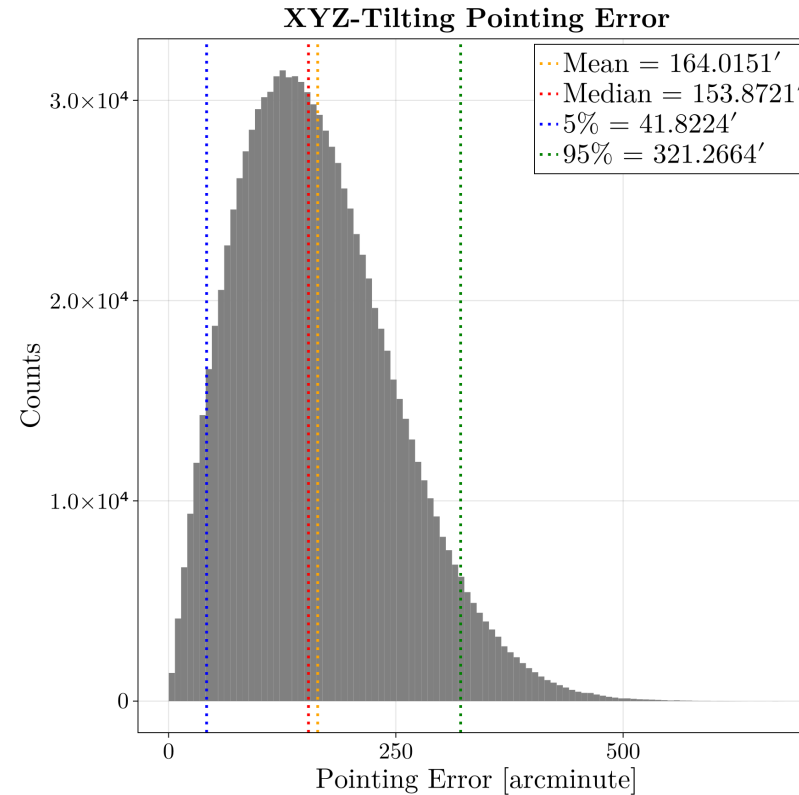
- We start the simulations with an **exaggerated value** of $\sigma = 1^\circ$ for debugging purposes, but we expect tilting errors approximately of the order of 10^2 arcsec.
- Later we studied how the **pointing error scales** with the variance σ^2 .

Ray-Tracing & MC Simulations

Multiple-Axis Single-Reflector Tilt

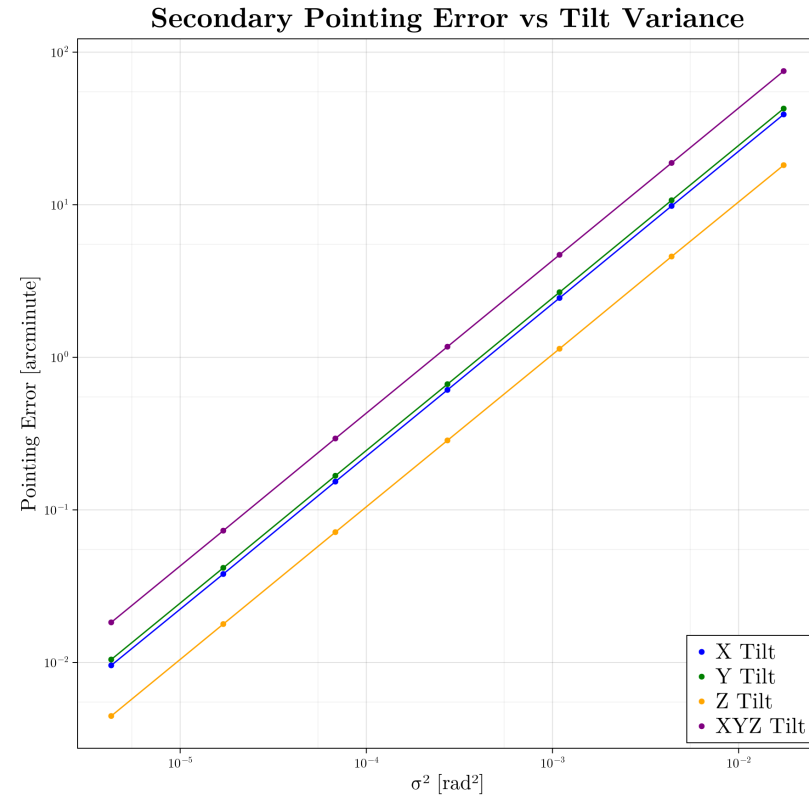
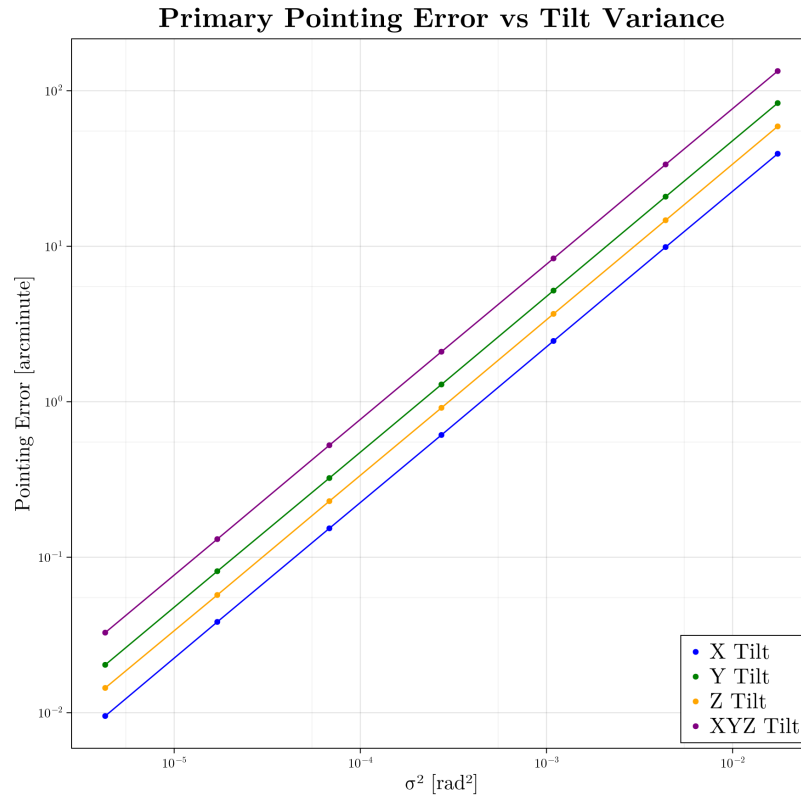


Multiple-Axis Multiple-Reflector Tilt



Ray-Tracing & MC Simulations

Pointing Error Trends



CMB Error Maps

Sky Observations using STRIPLINE.jl

We integrated the pointing direction evaluated by our ray-tracer into STRIPLINE.jl, the end-to-end scanning strategy pipeline for STRIP.

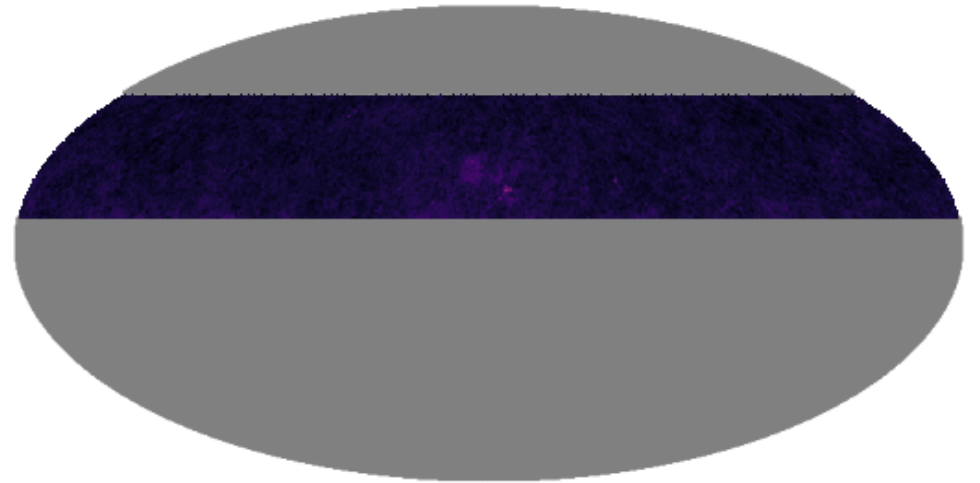


Figure 9: Simulated sky map observed by STRIP in a year of sky observation. The telescope, situated at the Canary Islands, will observe a strip of northern sky.

Sky Observations using STRIPLINE.jl

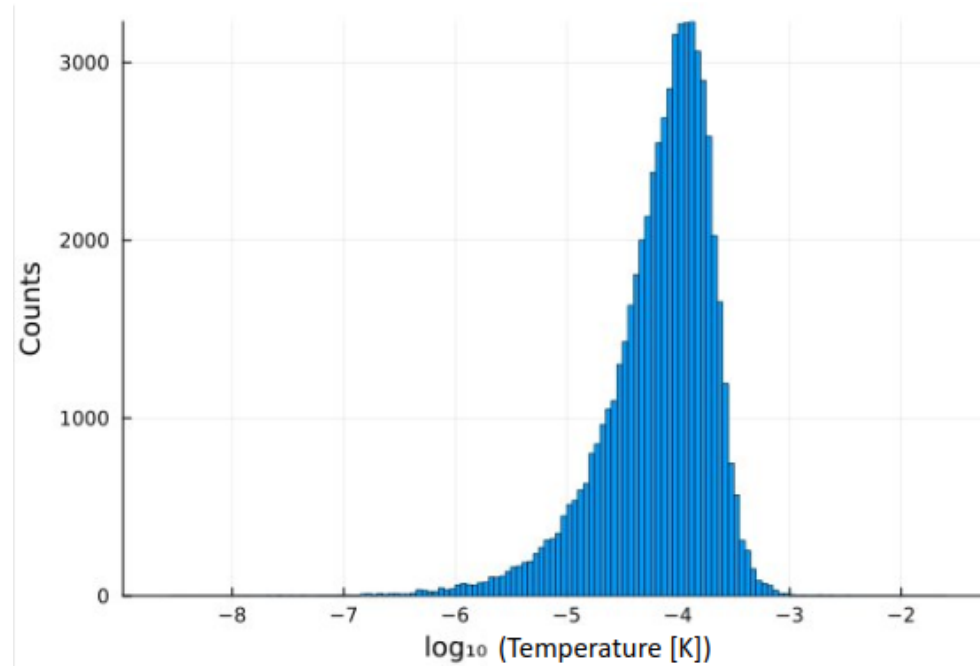


Figure 10: Histogram of the logarithm in base 10 of the error maps pixels.

Conclusion and Outlooks

Conclusion

- Through the **development of a ray-tracing engine** using the Julia programming language, the application with **Monte Carlo techniques** and the **integration into** the STRIP simulation **pipeline**, this study effectively demonstrates a flexible and computationally efficient **framework for optical validation** of CMB measurements.

Conclusion

- Through the **development of a ray-tracing engine** using the Julia programming language, the application with **Monte Carlo techniques** and the **integration into** the STRIP simulation **pipeline**, this study effectively demonstrates a flexible and computationally efficient **framework for optical validation** of CMB measurements.
- To our knowledge, this represents the **first integration** of a geometrical ray-tracing engine into a full end-to-end scanning strategy pipeline.

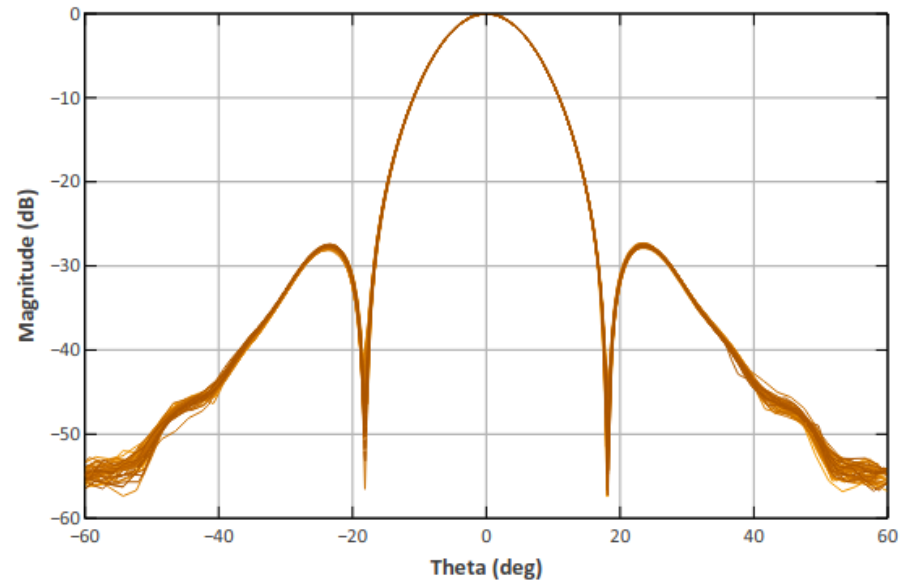
- Investigate **tilt angles correlations**.
- **AD** for uncertainty propagation, optimization problems and visualization of the pointing error gradients.
- Model surface imperfections with **Zernike Polynomials**.
- **Diffraction effects** with PO and MoM.
- Analyze **different kinds of perturbations** on **different telescopes**.

Thank you for the attention

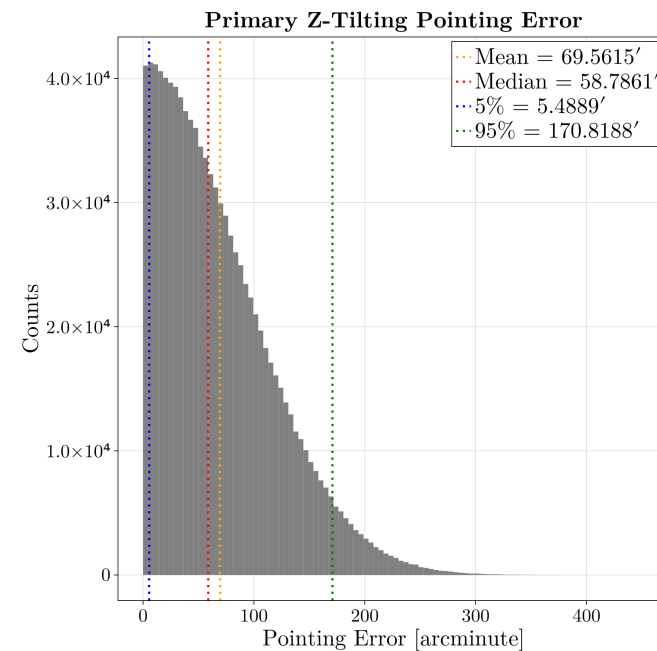
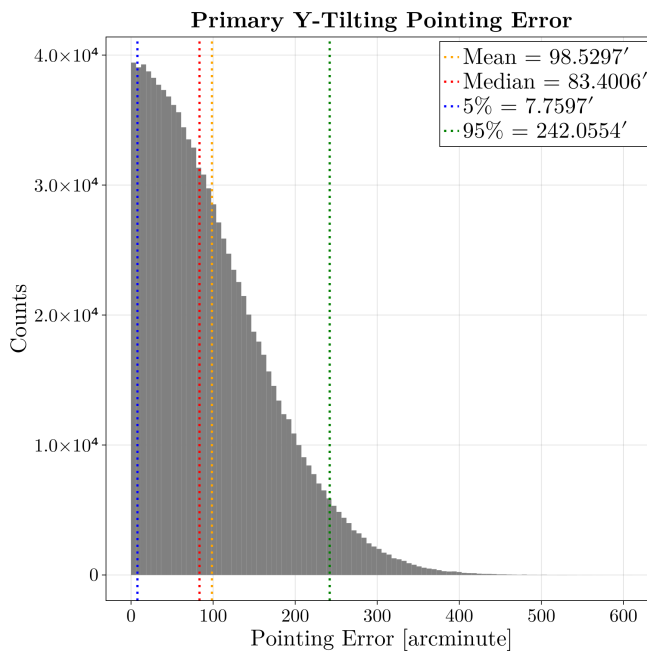
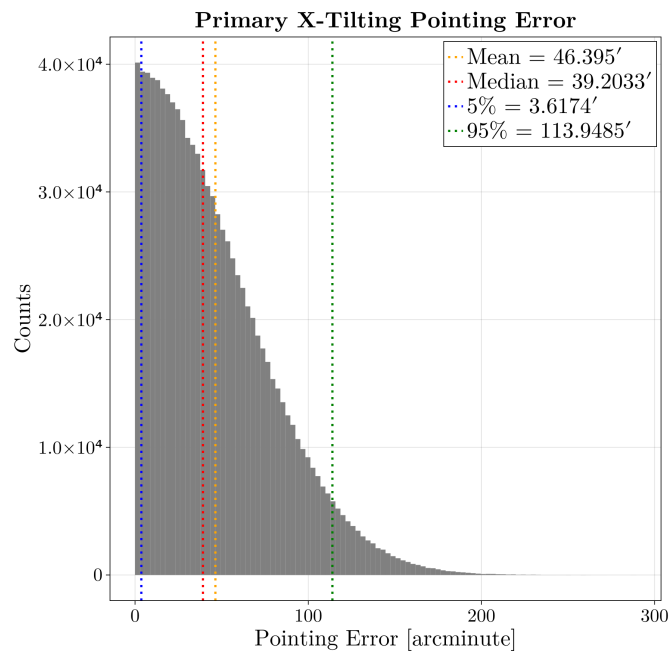
Appendix

Appendix

Feed's radiation pattern

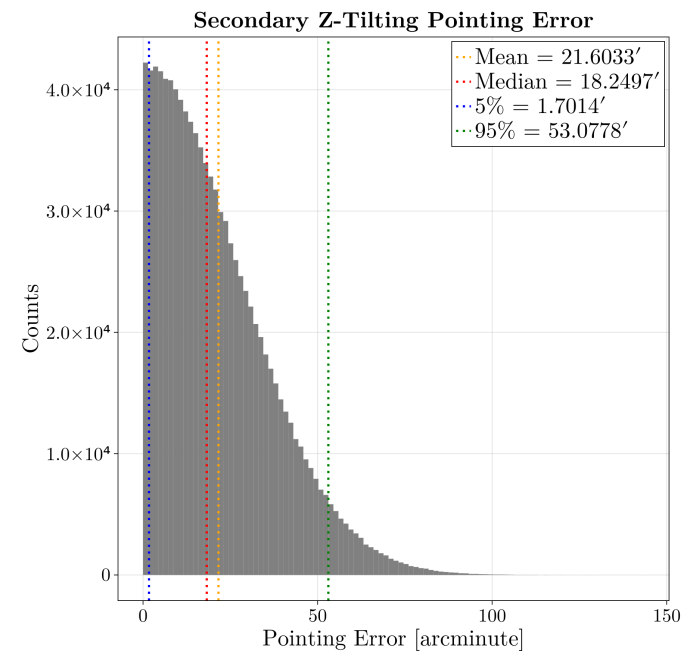
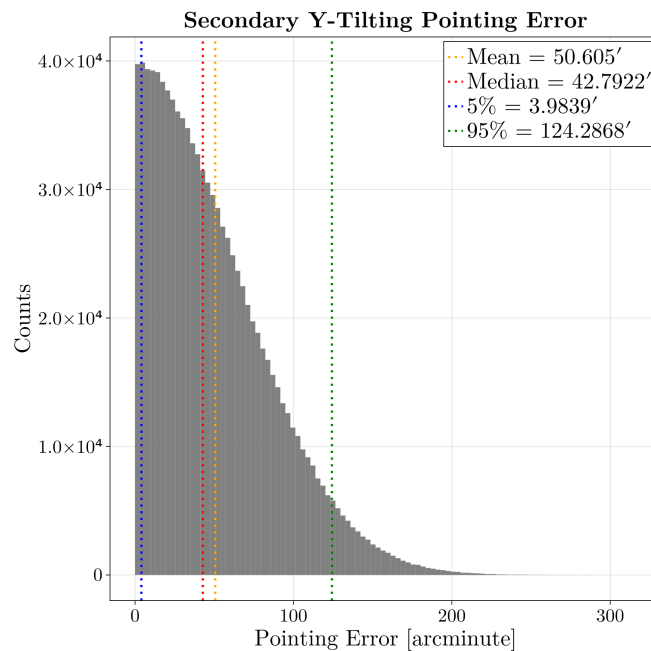
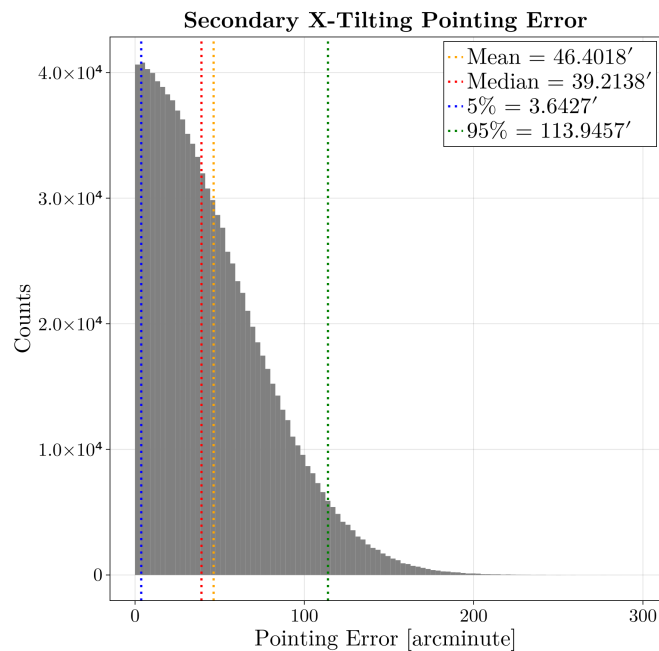


Single-Axis Single-Reflector Tilt



Appendix

Single-Axis Single-Reflector Tilt



Let's build a ray-tracing engine!

Julia Programming Language

- Modern language for **scientific** computing.
- Combines Python-like syntax with **C-like performance**.
- Excellent packages: `StaticArrays.jl`, `CoordinateTransformations.jl` and `Makie.jl`.
- Handles **units** and **automatic differentiation** very easily with type dispatch.
- An existing end-to-end simulation, `STRIPLINE.jl`, that models the **scanning strategy** of the telescope up to the creation of **frequency maps**.

Julia implementation of a Ray

```
using StaticArrays
```

Julia

```
struct Ray{0, D}
```

```
    origin::SVector{3, 0}
```

```
    direction::SVector{3, D}
```

```
end
```

```
(ray::Ray)(t::Real) = ray.origin + ray.direction * t
```

Julia implementation of an AbstractSurface

```
abstract type AbstractSurface end
```

Julia

```
function ( $\Sigma$ ::AbstractSurface)(u, v) = SVector(u, v, height( $\Sigma$ , u, v))
```

```
function height( $\Sigma$ ::AbstractSurface, u, v) end
```

```
function gradient( $\Sigma$ ::AbstractSurface, u, v) end
```

```
function normal( $\Sigma$ ::AbstractSurface, u, v) end
```

```
function quad_intersection( $\Sigma$ ::AbstractSurface, ray) end
```

Julia implementation of a Reflector

```
using CoordinateTransformations
```

Julia

```
const Mapping{T} = Tuple{T, Transformation}
```

```
struct Reflector{ $\Sigma$  <: AbstractSurface,  $\Gamma$  <: AbstractRim}
```

```
    surface::Mapping{ $\Sigma$ }
```

```
    rim::Mapping{ $\Gamma$ }
```

```
    name::String
```

```
end
```